

Lecture 5 - Ideal Gas

Last time:

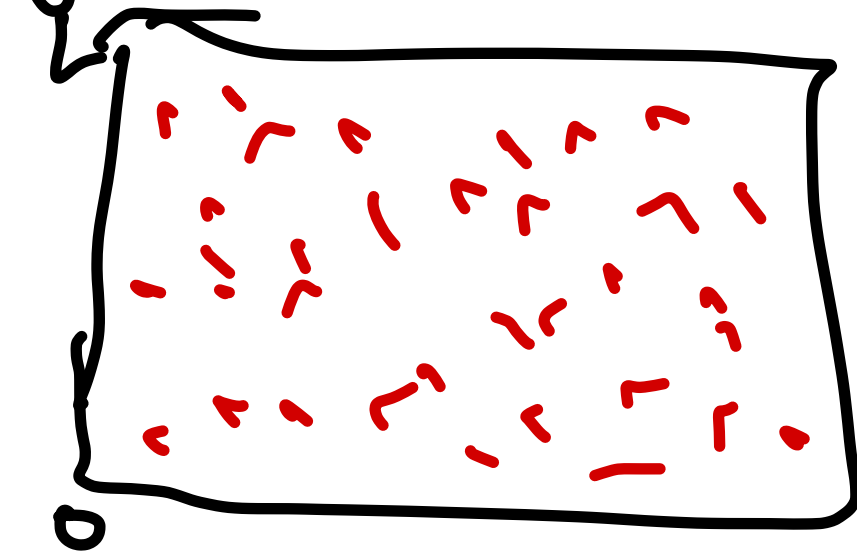
Entropy: state function

- doesn't decrease
(on average)

For an isolated system

$$\frac{1}{T} = \left(\frac{\partial S}{\partial E} \right)_{N, V} \quad - \quad P = \left(\frac{\partial S}{\partial V} \right)_{N, E}$$

Ideal gas:



const

N, V, E

$$V = L^d$$

density: $\rho = N/V$ low

molecules don't "feel" each other

$$U(x) = \begin{cases} 0 & \text{if } x_i \in (0, L) \quad i=1 \dots 3N \\ \infty & \text{otherwise} \end{cases}$$

$$S = k_B \ln \Omega(N, V, E)$$

$$\Omega(N, V, E) = \frac{E_0}{h^{3N} N!} \int \delta(H(\vec{x}) - E) d\vec{x}$$

$$d\vec{x} = dx_1 dx_2 \dots dx_{3N} dp_1 dp_2 \dots dp_{3N}$$

$$\int_{-\infty}^{\infty} dx_1 \dots \int_{-\infty}^{\infty} dp$$

$$H(\vec{x}) = \sum_{i=1}^{3N} \frac{p_i^2}{2m}$$

prefactor $\frac{\epsilon_0}{h^{3N} N!}$

• $\epsilon_0!$ $\delta(x)$ has units of $\frac{1}{[x]}$

$$\int_{-\infty}^{\infty} dx \delta(x) = 1$$

eg $\int_a^b dx \cdot 2 = 2 \cdot (b-a)$

$$S = k_B \log \Omega = k_B \log \epsilon_0 + k_B \log(\quad)$$

• $N!$

• a

• b

• c

1 a

b

c

2 | a b c
| a b c
| a b c

3 | a b c
| a b c
| a b c
| a b c

[$N_A! N_B! \dots$]

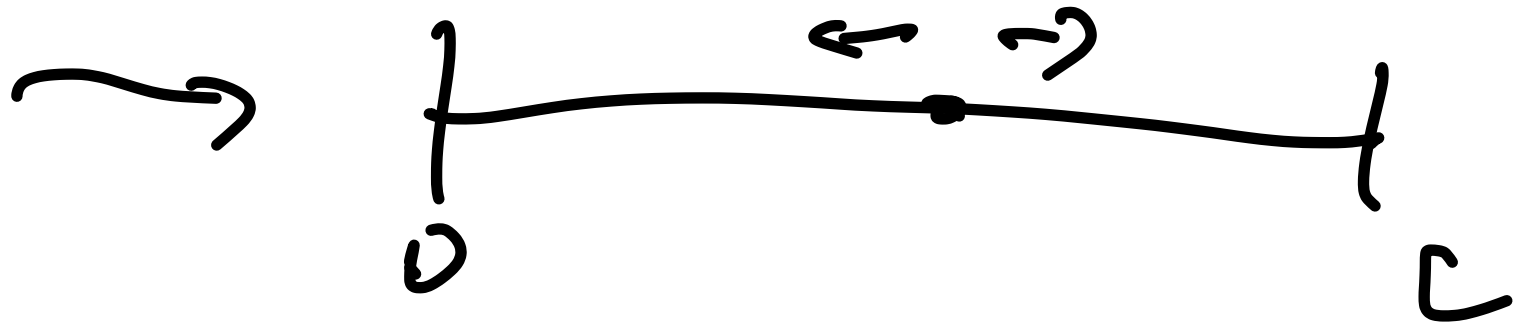
how many ways can we label these molecules?

$$N \cdot (N-1) \cdot (N-2) \dots 1$$

• h^{3N}

$$\int dx_1 \dots \int dp_1 \dots dp_{3N}$$

x.p



how many states does this have

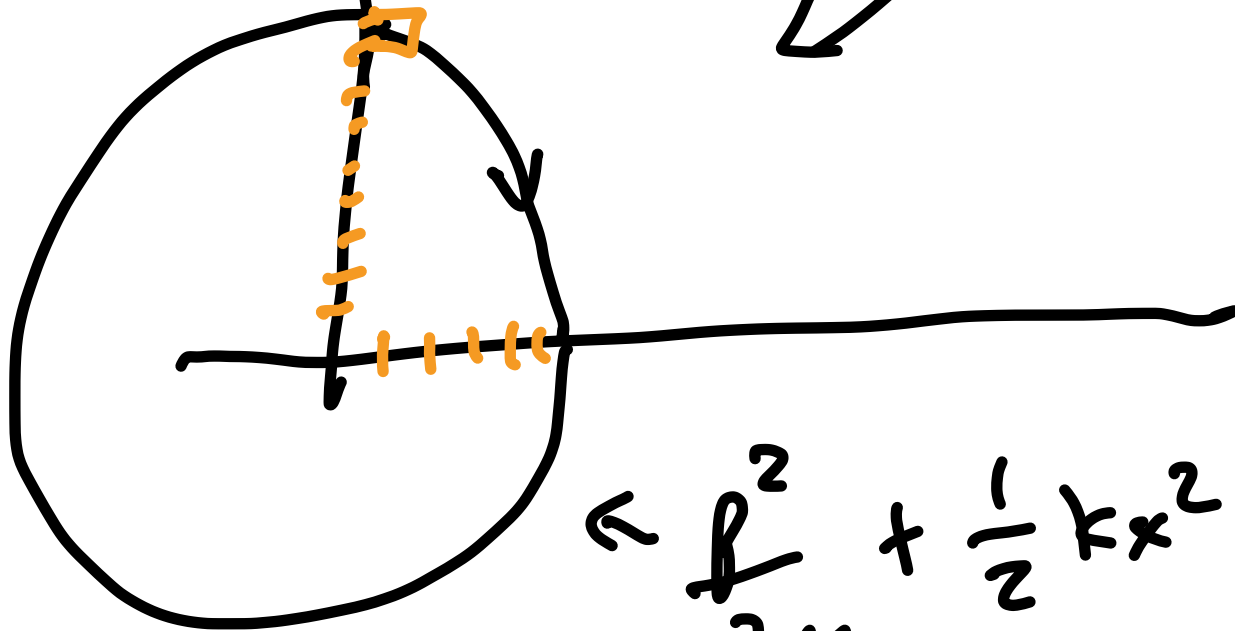
p

$$p^2/2m = E$$

$$V(x) = E$$

can't know x_i, p_i simultaneously

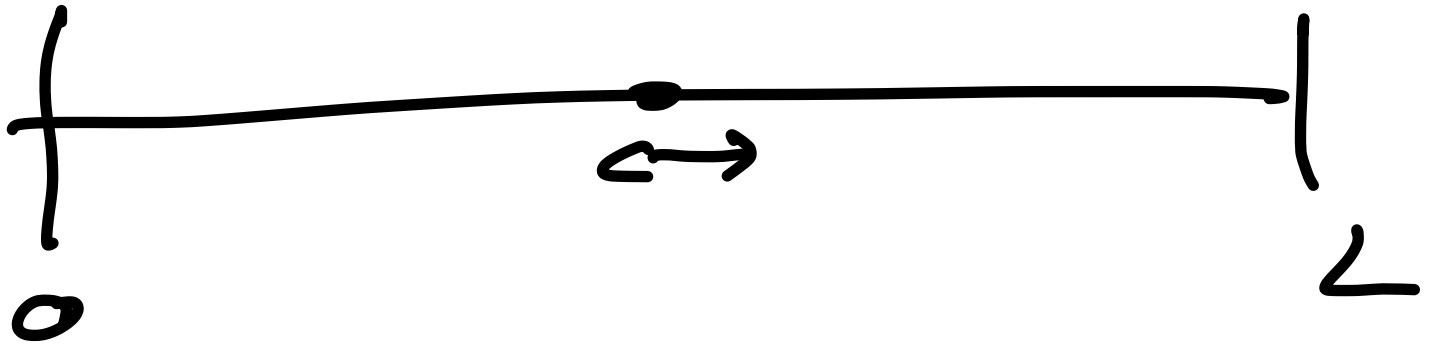
$$\Delta x \Delta p \geq \frac{\hbar}{2}$$



$$p^2/2m + \frac{1}{2} kx^2 = E$$

$x \cdot p \approx \hbar$

Ω for an ideal gas
 consider 1 particle in 1D



$$\Omega = \frac{E_0}{h \cdot (1!)} \int_0^L dx \int_{-\infty}^{\infty} dp \delta\left(\frac{p^2}{2m} - \varepsilon\right)$$

$\underbrace{\hspace{10em}}_{\mathcal{H}(\vec{x})}$

$$y = \sqrt{2m} p \quad dy = \sqrt{2m} dp$$

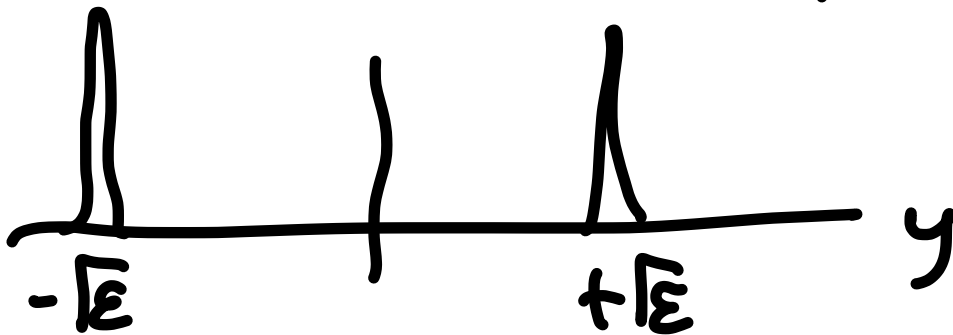
$$= \frac{E_0}{h} \cdot \sqrt{2m} L \cdot \int_{-\infty}^{\infty} dy \delta(y^2 - \varepsilon)$$

$$\delta(y^2 - \epsilon) = \delta((y - \sqrt{\epsilon})(y + \sqrt{\epsilon}))$$

$$\rightarrow = \frac{1}{2\sqrt{\epsilon}} [\delta(y - \sqrt{\epsilon}) + \delta(y + \sqrt{\epsilon})]$$

$$[\delta(F(x)) = \sum_{\text{roots}} \delta(\dots)]$$

$$\Omega_{\text{Id}} = \frac{2\epsilon_0 \sqrt{2m}}{h} \int_{-\infty}^{\infty} dy \cdot \frac{1}{2\sqrt{\epsilon}} \cdot [\delta(y - \sqrt{\epsilon}) + \delta(y + \sqrt{\epsilon})]$$



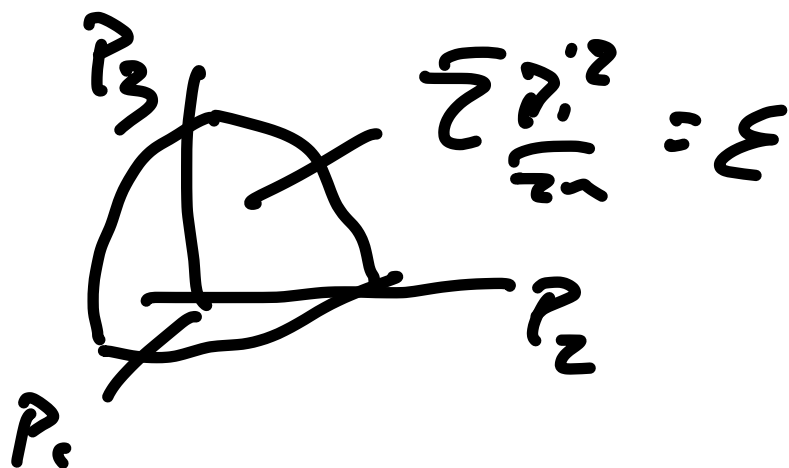
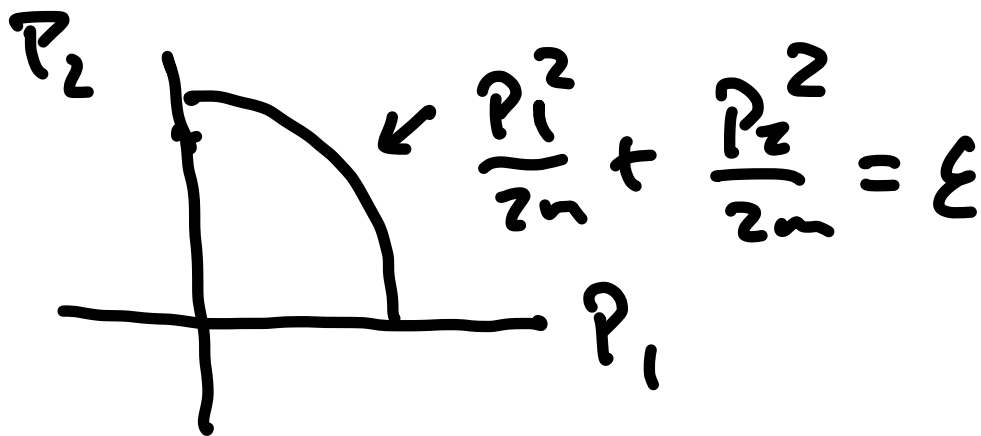
$$= \frac{2\epsilon_0 \sqrt{2m}}{h \sqrt{\epsilon}}$$

$$\Omega(N, U, \epsilon) = \frac{\epsilon_0}{h^{3N} N!} \int_0^L dx_1 \dots \int_{-\infty}^{\infty} dp_1 \dots dp_{3N}$$

$$L^{3N} = V^N$$

$$\delta\left(\sum_{i=1}^{3N} \frac{p_i^2}{2m} - \epsilon\right)$$

$$= \frac{\epsilon_0 V^N}{h^{3N} N!} \int dp^{3N} \delta\left(\sum_{i=1}^{3N} \frac{p_i^2}{2m} - \epsilon\right)$$



$$x^2 + y^2 + z^2 = r^2$$

$$\int d\theta d\varphi$$

$$dx dy dz = r^2 \sin\theta \underbrace{dr d\theta d\varphi}$$

$$\underbrace{4\pi r^2 dr}$$

Surface area of the
sphere

$$p_1^2 + p_2^2 + \dots + p_{3N}^2 = r^2$$



$$dp_1 dp_2 \dots dp_{3N} = \underbrace{\int_{3N-1}} r^{3N-1} dr$$

$$\Omega \stackrel{\text{large } N}{\approx} \frac{\epsilon_0}{N!} \left(\frac{V}{h^3} \left(\frac{4\pi m \epsilon_0}{3N} \right)^{3/2} \right)^N$$

↑ depends on N, V, ϵ

$$S = k_B \ln \Omega(N, V, \epsilon)$$

$$\frac{1}{T} = \left(\frac{\partial S}{\partial \epsilon} \right)_{N, V} = k_B \left(\frac{\partial \ln \Omega}{\partial \epsilon} \right)_{N, V}$$

$$\begin{aligned} \ln \Omega &= \ln \left[\epsilon^{3/2 N} \cdot \text{stuff} \right] \\ &= \frac{3}{2} N \ln \epsilon + \ln(\text{stuff}) \end{aligned}$$

$$\frac{1}{T} = k_B \cdot \frac{3}{2} N \cdot \frac{1}{\epsilon} \Rightarrow \epsilon = \frac{3}{2} N k_B T \quad \checkmark$$

$$\frac{P}{T} = \left(\frac{\partial S}{\partial U} \right)_{N, E}$$

$$\ln S = \ln V^N + \ln(\text{stuff})$$

$N \ln V$

$$\frac{P}{T} = k_B \frac{\partial (N \ln V)}{\partial V} + C$$

$$\frac{P}{T} = k_B \frac{N}{V} \Rightarrow PV = Nk_B T \checkmark$$

$= nRT$

$$\Omega = \frac{\epsilon_0}{N!} \left(\frac{V}{h^3} \left(\frac{4\pi m \epsilon e}{3N} \right)^{3/2} \right)^N$$

$\epsilon = \frac{3}{2} N k_B T$

$$= \frac{\epsilon_0}{N!} \left(\frac{V}{h^3} \left(2\pi m k_B T \right)^{3/2} \right)^N e^{3N/2}$$

$$= \frac{\epsilon_0}{N!} \left(V \cdot \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \right)^N e^{3N/2}$$

$$\Lambda = \frac{h^2}{\sqrt{2\pi m k_B T}}$$

$$= \frac{\epsilon_0}{N!} \left(\frac{V}{\Lambda^3} \right)^N e^{3N/2}$$

$$S = k_B \ln \Omega$$

$$\Omega = \frac{\epsilon_0}{N!} \left(\frac{V}{\lambda^3} \right)^N e^{3N/2}$$

$$= N k_B \ln \left(\frac{V}{\lambda^3} \right) - k_B \ln(N!) + \frac{3}{2} N k_B + \text{const}$$

$$\ln(N!) \approx N \ln N - N$$

$$= N k_B \ln \left(\frac{V}{\lambda^3 N} \right) + \frac{5}{2} N k_B + \dots$$

$$= \left\{ N k_B \ln \left[\frac{V}{N} \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \right] + \frac{5}{2} N k_B \right\} + \dots$$

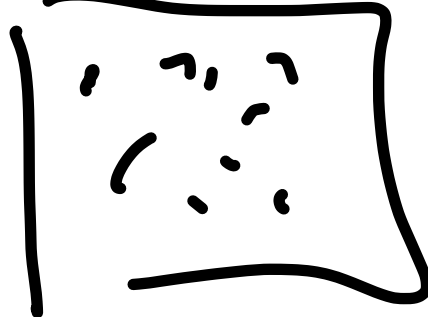
Sackur - Tetrode

why do we need $N!$

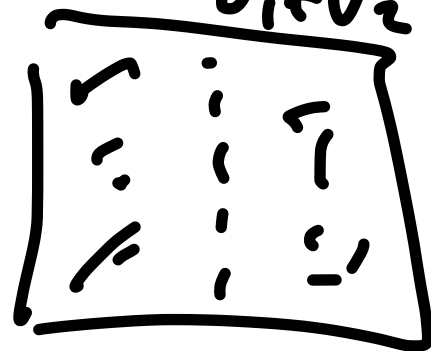
N_1, V_1, E_1



N_2, V_2, E_2



$N_1 + N_2$
 $V_1 + V_2$



Gibbs' Paradox

HW: $S_{\text{combined}} = S_1 + S_2$