

Lecture 25

Intro to Quantum Stat
Mech, pt 2

Ideal Gas Distinguishable Particles (Neglect spin)

$$Q(N, V, T) = q^N = \left(\sum_n e^{-\beta \epsilon_n} \right)^N \quad \text{"Boltzmann particles"}$$

v/ Occupation #s,

$$Q(N, V, T) = \sum_{\{f\}} g(\{f\}) e^{-\beta \sum_n \epsilon_n f_n}$$

$$g(\{f\}) = \frac{N!}{\prod_n f_n!}$$

← new state, distinguishable
same occupation #s

eg, 2 states, $f = f_1, f_2$ $g(f_1, f_2) = \frac{N!}{f_1! f_2!} = \frac{N!}{f_1! (N-f_1)!}$

$$Q(N, V, T) = \sum_{\{f\}} \frac{N!}{\prod_n f_n!} \prod_n e^{-\beta f_n \epsilon_n} = \left(\sum_n e^{-\beta \epsilon_n} \right)^N \quad !!$$

multinomial expansion

So, 1 particle $\sum_n e^{-\beta \epsilon_n} = \sum_n e^{-2\pi^2 \beta \hbar^2 n^2 / m L^2}$

when $L \rightarrow \infty$, spacing becomes continuous, sum

$$\Rightarrow \int dn e^{-2\pi^2 \beta \hbar^2 n^2 / m L^2} = 4\pi \int_0^\infty dn n^2 e^{-2\pi^2 \beta \hbar^2 n^2 / m L^2}$$

result $= V \left(\frac{m}{2\pi\beta\hbar^2} \right)^{3/2} = \frac{V}{\Lambda^3}$ ← same as classical ideal gas

All QM must be in spin interactions

Identical fermions (bosons)

Exchanging particles is same physical state, so

$$g(\{f_{\vec{n}m}\}) = 1, \text{ but}$$

$$f_{\vec{n}m} = 0, 1 \text{ for fermions}$$

$$Q(N, U, T) = \sum_{\{f_{\vec{n}m}\}} e^{-\beta \sum_m \sum_n f_{nm} E_n} = \sum_{\{f_{\vec{n}m}\}} \prod_n \prod_m e^{-\beta f_{nm} E_n}$$

w/ restriction $\sum_m \sum_n f_{nm} = N$

which is hard to keep track of

Instead better to use $Z(\mu, U, T) = \sum_{N=0}^{\infty} \zeta^N Q(N, U, T)$

$$= \sum_{N=0}^{\infty} e^{\beta \mu N} \sum_{\{f_{nm}\}} \prod_m \prod_n e^{-\beta f_{nm} E_n}$$

Same results in therm limit

still subject to occupation #, but outer sum

relieves restriction as all things will be added

$$\text{ie } Z(\mu, U, T) = \sum_{\{f_{nm}\}} \prod_m \prod_n e^{+\beta(\mu - E_n) f_{nm}}$$

has all values of N taken into account

$$\text{B/c } \sum_{f_1} \sum_{f_2} \sum_{f_3} \dots e^{+\beta(\mu - E_1) f_1} e^{+\beta(\mu - E_2) f_2} \dots = \left(\sum_{f_1} e^{+\beta(\mu - E_1) f_1} \sum_{f_2} e^{+\beta(\mu - E_2) f_2} \dots \right)$$

$$\Rightarrow Z = \prod_m \prod_n \sum_{\{f_{nm}\}} e^{\beta(\mu - \epsilon_n) f_{nm}}$$

for fermions $f_{nm} = 0$ or 1

$$\star Z = \prod_m \prod_n (1 + e^{\beta(\mu - \epsilon_n)}) \quad (\text{fermions})$$

For bosons, $f_{nm} = 0 \rightarrow \infty$, like last time

$$Z = \prod_m \prod_n \frac{1}{1 - e^{\beta(\mu - \epsilon_n)}} \quad (\text{bosons})$$

Product over $m = 0 \dots 2s+1 = g$

$$Z = \prod_n (1 + a e^{\beta(\mu - \epsilon_n)})^{g \cdot a}$$

$a = 1$ for fermions, -1 for bosons

$$PV/k_B T = \ln Z = g a \sum_n \ln(1 + a e^{\beta(\mu - \epsilon_n)})$$

$$= g a \sum_n \ln(1 + a \underbrace{\xi e^{-\beta \epsilon_n}}_{\substack{\leftarrow e^{\beta \mu}, \\ \text{fugacity}}})$$

$$\langle N \rangle = \xi \frac{\partial}{\partial \xi} \ln Z = \xi \sum_n \frac{a e^{-\beta \epsilon_n}}{1 + a \xi e^{-\beta \epsilon_n}} = g \sum_n \frac{e^{\beta(\mu - \epsilon_n)}}{1 + a e^{\beta(\mu - \epsilon_n)}}$$

Now, comparing to a fixed N situation

$$N = \sum_m \sum_n f_{nm}$$

$$\Rightarrow \langle N \rangle = \sum_m \sum_n \langle f_{nm} \rangle$$

$$\Rightarrow \text{we can associate } \langle f_{nm} \rangle = \frac{e^{\beta(\mu - \epsilon_n)}}{1 + e^{\beta(\mu - \epsilon_n)}} \text{ for fermions}$$

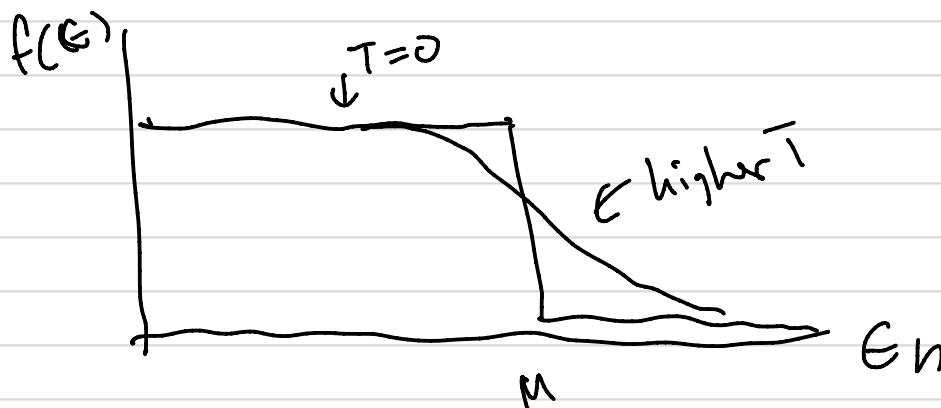
$$\Rightarrow \frac{1}{1 + e^{-\beta(\mu - \epsilon_n)}} = \frac{1}{1 + e^{\beta(\epsilon_n - \mu)}} \quad (\text{fermi-dirac distribution})$$

consider low temp, high β ,

$$\text{for } \epsilon_n - \mu > 0, \quad \langle f \rangle \rightarrow 0$$

$$\epsilon_n - \mu < 0 \quad \langle f \rangle \rightarrow 1$$

$$\langle f_{nm} \rangle^{T=0} = \begin{cases} 0 & \epsilon_n > \mu \\ 1 & \epsilon_n < \mu \end{cases} = \Theta(\mu - \epsilon_n)$$



The energy where states above are unoccupied is the "fermi energy"

Ideal boson gas

Don't have time to go into details, but

boson gas can undergo Bose-Einstein condensation

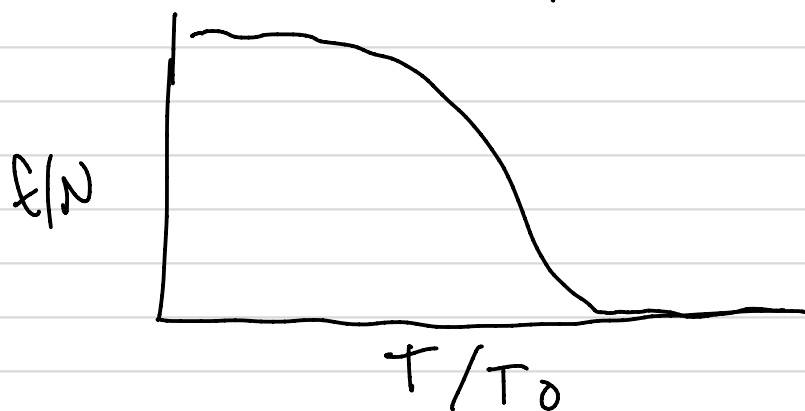
where at low T all particles go

into the same state

$$\langle f_{nm} \rangle = \frac{1}{e^{\beta(\epsilon_n - \mu)} - 1}$$

$$\langle f_{0m} \rangle = \frac{e^{\beta\mu}}{1 - e^{\beta\mu}}$$

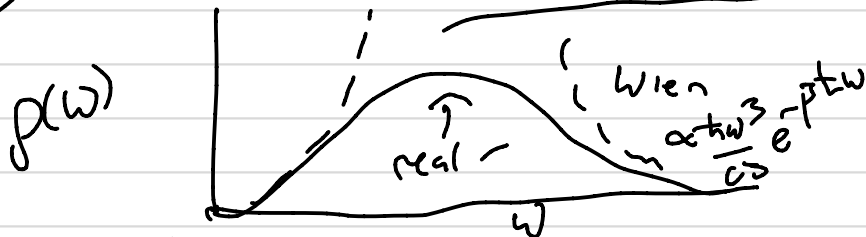
at same temp, can show



Black-body radiation



← heat radiating object, will absorb & emit radiation



Rayleigh -
Jeans
 $\propto \frac{\omega^2 k_B T}{c^2}$
(only good small ω)

Photons have spin 1, bosons

also emitting & absorbing photons, so just like our grand canonical treatment

Only standing waves commensurate w/ the size of the cavity are allowed, kind of like particle in a box

Hence only $\vec{k} = \frac{\pi}{L} \vec{n}$ $n_x, n_y, n_z = 1, 2, \dots$

$E = \hbar c |\vec{k}| = \hbar c k$
momentum = $\hbar \vec{k} = \hbar k$

$k^2 = \pi^2/L^2 (n_x^2 + n_y^2 + n_z^2)$

Number of waves w/ momentum $< |k|$ is volume of sphere upper quadrant (all pos n_i)
 $\phi(k) = \frac{4}{3} \pi R^3 / 8 = \frac{1}{6} \pi \left(\frac{Lk}{\pi} \right)^3 = V k^3 / 6\pi^2$

btwn k & dk
 $\omega(k) dk = \frac{d\phi}{dk} dk = \frac{V k^2}{2\pi^2} dk$

$\times 2$ for 2 polarizations

One can show (no time here)
that $\mu = 0$ for an ideal gas of photons

In this case we know

$$\langle f_\epsilon \rangle = \frac{1}{e^{\beta\epsilon} - 1}$$

$$\Rightarrow \langle \mathcal{E} \rangle = \sum_{\epsilon} \epsilon f_\epsilon = \sum_{\mathbf{k}} \frac{\hbar c k}{e^{\beta \hbar c k} - 1}$$

if \mathbf{k} 's close together, can convert to
integral w/ DOS

$$\langle \mathcal{E} \rangle \rightarrow \int_0^\infty dk \omega(k) \frac{\hbar c k}{e^{\beta \hbar c k} - 1}$$

$$= \frac{V \hbar c}{\pi^2} \int_0^\infty dk \frac{k^3}{e^{\beta \hbar c k} - 1} \quad \omega = ck$$

$$= \frac{V \hbar}{\pi^2} \int_0^\infty d\omega \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

$$\epsilon/v \equiv \int_0^\infty \rho(\omega, T) d\omega$$

$$\Rightarrow \rho(\omega, T) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

← Planck
formula

$$\text{@ low } \omega, e^{\beta \hbar \omega} \approx 1 + \beta \hbar \omega$$

$$\Rightarrow \rho(\hbar \omega < c k_B T, T) \approx \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{\beta \hbar \omega}$$

$$\propto \frac{\omega^2 k_B T}{c^3}$$

$$\& \text{@ high freq } \rho(\omega, T) \propto \frac{\hbar \omega^3}{c^3} e^{-\beta \hbar \omega}$$

Summary: in stat mech, we consider averages over many particles & it gives us powerful tools & can derive sim methods to solve problems where exact solns seem impossible!

- We also connect microscopic quantities to bulk observable measurements
- but more importantly can predict fluxes in small # experiments & heterogeneous environments
- and may be even out of eq
- Hopefully, you can apply these ideas in your own work!